

# THE EVENING NEWS.

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## UNIVERSITY LECTURES.

### ASTRONOMY.

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#### LECTURE IV.

**The Seasons.**—Subdivisions of the Equinoxes.—Precession of the Equinoxes.—Alderal and Tropical year.—Revolution of the elliptic orbit.—Anomalous year.—Geocentric and Heliocentric places of a Heavenly Body.—Earth's Mean and True Longitude.—Mean and Apparent time.—Distributions of Temperature.—Hottest and Coldest days of the year.—Permanency of the mean annual temperature.—Invariability of the earth's diurnal and annual Periods.—Stability of the Laws of Motion.—Interior Temperature of the earth as affected by the Sun.—Ocean Temperature.—Cause of the great Currents in the Ocean.—Atmospheric Phenomena.—Reflections on the Origin of Planetary Motion.

This first subject which we propose to investigate in this lecture is the Seasons. During the time that the earth performs one annual revolution, the inhabitants experience a variety of seasons.

Those who live in the southern hemisphere have their seasons in the reverse order of those in the northern. December, January and February are their summer months; while here, they are our winter months. Their spring corresponds to our autumn; their winter to our summer; their autumn to our spring. When the days in the northern hemisphere are the longest, in the southern they are the shortest; and vice versa; when they are the shortest, they are the longest there. From the 21st of March to the 21st of September, the sun shines without any intermission on our north pole, while the south pole during that time is enveloped in darkness. From the 21st of September to the 21st of March, the south pole is constantly enlightened by the sun, while our north pole is left in darkness. The whole order of the seasons in the northern hemisphere is repeated in the southern, but during the opposite time of year.

If the earth revolved around the sun directly from west to east, that is, if the plane of the earth's orbit coincided with the plane of the equator, there would be no variety of seasons; and also the days and nights over the whole earth would be of equal length. If the earth revolved around the sun from south to north, and back again to the south, then our seasons would have the greatest possible change that could be given to them. The difference between the length of days and nights would increase with much greater rapidity, and the extremes of temperature, between summer and winter, would also be far greater. On about the 10th of March, the days would be about the 10th of May, the days would increase from 12 hours to 24.

All the nights would decrease from 12 to nothing. From the 10th of May to the 2nd of August, the sun would not set at night he would be seen among our circumpolar stars, exhibiting the same apparent phenomena manifested by those stars. About the 2nd of August night would again set in, the length of which would now increase until the 21st of September, when the days and nights would again be equal. From the 21st of September the length of the nights would increase until about the 11th of November, when the sun would set and remain below the southern horizon about eighty days, or until about the 31st of January, when the day would be seen in being only a few minutes long at first, but increasing rapidly in length until the 21st of March, when day and night would again be equal.

Thus if the earth revolved in an orbit whose plane was perpendicular to the plane of the equator, the vicissitudes of the seasons and the variations of day and night would be such as to render our globe unfit for the habitation of man. At one season of the year he would be scorched not only with a vertical sun, but with an accumulation of heat arising from the great length of the day; while in another season he would be exposed to all the severity of cold experienced in the polar regions.

If the earth should revolve around the sun in any other direction, except the two that we have just considered, the length of the seasons, and of day and night would be proportional to the inclination of the elliptic to the plane of the equator, as the angle of inclination increases so would the difference between the length of the day and night. The inclination of the two planes is called the obliquity of the elliptic, which is about 23° 56' 30".

We shall next explain what is meant by the Geocentric and Heliocentric places of a heavenly body.

The Geocentric place is its position as it would be seen from the center of the earth.

The Heliocentric place is its position as it would be seen from the center of the sun.

The mean place is the position as it would be seen from the center of the sun.

The true longitude is the mean place of the sun in its elliptic orbit.

The mean longitude is the mean place of the sun in its elliptic orbit.

The anomalous year is the mean place of the sun in its elliptic orbit.

The tropical year is the mean place of the sun in its elliptic orbit.

The sidereal year is the mean place of the sun in its elliptic orbit.

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