fireritation and Contri Juget Forces, br. **ORSON PRATT, Sen.**, Delivered before the Members of the Normal Institute, at the Descret University, on Wednesday Even

LECTURE

GRAVITATING FORCE. - Its Tw: Manifestations-Failing Bodies-Universal Tendency of Worlds to fall to cach other Three Laws of Gravitating Force - When Dis-CENTRIFCOAL FORCE Antagonistic to Gravity-Unic: Fendency of Worlds to Move : Straight Lines, instead of Curred Orbits-Curvillinear Orbits in Fronte of Two Forces-Four Onter angenible viz, the Circle, Ellipse, Marabola and Hyperbola - U. tr what Circumstances each is produc-ed-Measure of Projected Forces -Illustrated by Whirling Weights -Revolving Worlds-Law of Distance and Velocity-Weight of Worlds Colestial Mechanism an Exhibition of Infinite Wisdom

## and Power.

1.-GRAVITATION is a force exeri-ed between all particles of matter in the universe. Two or more par-ticles of masses of matter, placed in contact, have a mutual tendency to press towards each other's centers. This pressure, however great may be the difference in the quan-tities of master, is equal and in op-posite directions; for instance, one pound weight exerts the same pressure, upon the surface of the earth, as the earth exerts upon the one pound weight; and as the two forces nitess in directions opposite to each other, the two bodies must re-

mutual pressures, then the bodies, compounded of these different kinds, would necessarily move with a uniformly accelerated motion in the direction of the greater pressure. This may be represented by two steam vessels of equal weight in contact, pressing in opposite direc-tions: if the force of steam on the two boats is equal, they will re-main at rest; if the force of one preponderates over the other, there will be, not only a pressure, but a motion in the direction the greater

thirty two-hundred thousandth part of circles an inch, in one second of time. While Nep-tune, being thirty times more distant from webeing the earth than the Sun, falls towards the

lem. All have beheld the moon circling we itself around the earth, at the distance, in we round numbers of about 240,000 miles, or at at the distance from the earth's centre, of 60 semi-diameters of the earth. All might have known that a body must moveln a straight line, when once put is motion, unless deflected from that straight line by come force. All might have income that the orbit or path of the moon, was curved a

that ever was unfolded to the mind of man. Thus we see that the same force which makes a stone fall LG inches in one see ond, causes the moon to fall 3,000 times less, during the same time. If gravitation did not vary in its intensity with the dis-tance, then the moon would fall with the same velocity as a stone near the earth's surface. If the velocities decrease merely as the distance, then the moon, being 60 semi-diameters of the carth distant, would fall 60 times slower, bat its falling relocity is known to be 60 times 60, or 3,600 times slower, being, as before stated, only about the one twentieth of an inch in a second. slower, being, as before stated, only about the one-twentieth of an inch in a second." The mean semi-diameter of the earth h about 3.956.9 miles: the mean distance of the Sun, according to the new value is about 91,430,000 miles. The latter dis-tance, divided by the former, gives a quo-

main at rest. If one kind of matter exerted a greater force than another in their multiple of the background of the background of the source of the carth; this squared or multiplied into itself is equal to 554,009,832. If the Sun containis equal to Sid, 059, 852. If the Sun contain-ed only the same amount of matter as the earth, then the latter would be deflected from the tangent of 1's orbit into a curve towards the Sun with a velocity 534,009,-833 times less per second than that of a failing body near the earth's surface; but failing body near the earth's surface; but observation proves that the earth falls towards the Sun with a velocity several hundred thousand times greater than it would on the above apposition, its devi-ation from the tangent of its orbit being about the one-eighth of an inch per see-ond; or more accurately, the earth falls towards the Sun about 1,607 times slower than a stone falls near the surface of the earth. Now if 534,009,832 be divided by 1.607 the onoient-will, be about 314.760

motion in the direction the greater force is exerted. 2.—Particles and masses of mat-ter, net only exert the force of pres-are when in contact, but when separated they manifest a mutual gravitation towards each other. By virtue of this force, when not re-strailed, they will move towards each other with a uniformly accel-erated motion. If a mass of mat-ter, elevated above the surface of the earth, is at liberty to move, it will fall towards the sur-face, while the earth, at liberty to move, it will fall towards the sur-face, while the earth, at liberty to move, it will fall towards the sur-face, while the earth, at liberty to move, it will fall towards the surface of the earth, at liberty to move, it will fall towards the surface of the distance, its falling velocity of the earth, at liberty to move, it will fall towards the surface of the distance, its falling velocity of the distance, its falling velocity the velocities of the two moving bodies will be inversely proportion-

circles around central forces. It is re-quired to determine what effect different velocities and distances will have in rela-

the earth than the Sun, falls towards the tarth, in one second, 500 times has than that of the Sun, or the two-thomsed inch.
Berhaps this midlenes may inquire, how this law of gravity is known to be true?
How is it known that the force called gravitation decreases as the square of the distance intreases? Who has ascended the heights of heaven to experiment upon distance increases? Who has a cended the heights of heaven to experiment upon this force? Who, at the vast distances of the moon, sub and planets, has observed any phenomena by which he could exactly, determine the fractions of an inch through which bodies would fail towards the earth in one second of time? The answer is that all mankind, of every age, have beheld the point of centrifugal force, be represented by the mone second of time?

Suppose the string, connecting the ball with the centre, to be varied in its length,

while is velocity remains the same, what will be the law of force exerted upon the string? Both experiment and calculation demonstrate that, . With the same velocity, the centrifugal fores varies inversity as the length of the string.

unless deflected from that straight line by come force. All might have igness a that the orbit or path of the moon, was curved towards the earth, being deflected every moment from a straight line, called its tangent, into a curved line. All might have known, by a simple calculation, the eract amount of this deviation from the tangent to the curve lines. All might the strangent is the extra line by reflecting that a body, revolving with the immortal Newton, some two centuries, ago, solved the simple problem. We say, simple problem, because it is now so deally that ever was unfolded to the mind of man.

When the length of the string remains the

same, the contribugal force varies diretly as the square of the velocity.

The square of the pelocity. Examples: A string one foot long, with a velocity equal to one generates a centrifugal force equal to one pound; if the same ball, at the end of the string, is made to revolve with twice the velocity, the force exerted upon the string will be four pounds: If the velocity is three times as great, the siring will be stretched nino pounds: If the revolving velocity, is five times as great, the lension on the string will be twenty-five pounds; if the circular velocity of the balls four times

circular velocity of the ball is 600 times as

moon is only, the 1 straight part of that velocity necessary, if gravitation remained its ease in intensity and it is tances. There-fore, the square of master of the con-rifugal forces at the mean distance of the con-rifugal forces at the mean distance of the moon. But the gravitating force towards the sance of the distance; that is, it must be about 3,506 times less at the moon's dis-tance, than at the surface of the earth. Hence, the moon fails towards the earth of the same at all distances; that is, it must be about 3,506 times less at the moon's dis-tance, than at the surface of the earth. Hence, the moon fails towards the earth of the same at all distances; the moon's dis-tance, the moon fails towards the earth of the same at all distances. He moon's dis-tance the central force. He earth is 10 hearts, 56 minutes, 4.56375 seconds, in order to balance the central force. He must be period must be greatly increased; that is, 19 hours, 56 minutes, and 6.58275 seconds, or rather 30,966.58375 seconds must be multi-plied by 50.69155, which is the number of times the velocity is decreased; the product will be 2,300,591.54555 seconds in the moon's mean solar days, which is the rack period will be 2,300,591.54555 seconds in the moon's mean solar days, which is the rack period will be 2,300,591.54555 seconds in the moon's mean solar days, which is the rack period second for the solar to be statily increased; the solar balance the velocity is decreased; the product will be 2,300,591.54555 seconds in the moon's mean solar days, which is the rack period second period will be a solar to statily and solar days, which is the rack period second period will be a solar to statice is the solar to statice of the solar the solar to statice of the solar to statice

by this importer. Those same dol-lars, invested in home made cloths, would have seconded the noble efforts of far-seeing men who are using their financial ability to make this people self sustaining. HOME INTEREST. to 84 minutes and 42.7146 seconds); this there-fore, is the precise time in which a body would revolve around the earth, near fis surface at the equator, without failing. The amount of centrifugal force genera-ted can be calculated by making the length of the string equal to 39.923,637 feet, and the velocity of the revolving one-pound weight equal to the square root of the length of the string; that is, 4574,23622 times swifter than the ball with one foot radius. The numerical formula will be

Banews W Booth Mrs

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Green A C Gamble Fred Grinn G G

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(4574.23822) = 1 pound ; that is, the one

20023637 pound of centrifugal force upwards from the centre of the earth, just balances the one pound of gravitating force towards the

the centre of the earth, just mainces the concround of gravitating force towards the centre. The moon is 50.96135 times the equatorial ratius of the earth distant. Why does it not fail to the earth? We answer, it would fail, if it did not revolve around the earth with a velocity sufficient to gravitate a centrifugal force upwards, equal to that of its gravita-tion downwards. Under the influence of these two oppesing forces, it is constantly maintained at its orbit distance. Tf the intensity of gravitation were the same at all distances, the moon would fail towards the earth, from the tangent to the curve of its orbit, 16 feet and 1 mch, in one second of time. To balance this downward tendency, its velocity around the earth would have to be increased 65.96435 times gratter than its present mean velocity; that is, the moon would have to revolve around the earth in 10 hours, 56 minutes and 6.55375 seconds. This is calculated by taking the square root of the moon's dis-tance, which we have already stated to be 50.96135 times the earth's radius. The square root of 59.96435 is equal to 7.743655; this multiplies into 5683.7146 seconds, gives a product of 20066.55375 seconds, which is could to 10 hours, 56 minutes, and 6.58275 this multiplies into 5683.7146 seconds, gives a product of 20066.55375 seconds, gives a product of 20066.55375 seconds, minutes a product of 20066.55375 seconds. The seconds gives a product of 20066.55375 seconds. The seconds gives a product of 20066.55375 seconds gives a product of an earth's induces, and 6.58275 seconds. One pound of centrifugal force generated Curtis Mr Curtis Mrs. Grinnin G Flora Hunter J W Chelds Mrs. Hunter J W Chelds Mrs. Alice A manda SJohnson Miss Clawson Mrs. Alice A Johnson J H Curtis Thaleas Jensen Mrs Catroll L Gertrude Catroll L Ouleff Mrs B Cortis Juo Davidson D Dahlstedt C Davis Miss

equal to 10 hours, 56 minutes, and excer-seconds. One pound of centrifugal force generated at the distance of the moon, by the whirl ing of a one pound weight, would require a velocity, greater than when whirled mean the surface of the earth. The eract ve locity necessary, will be found by taking the square root of the moon's distance, o 7.743665 times the velocity of a revolving body at the earth's surface. Then w shall have the numerical formula, thus-

(7.743865) = 1 pound.

(7.743865) = 1 pound. 59.98435 But from observation we know that the moon's real velocity around the earth is not 7.743865 times swifter than a revolving body at the surface of the earth; indeed, it is known to have a mean velocity 7.743865 times slower than that of a body at the earth's surface, or 59.96455 times slower than the velocity necessary to generate 1 pound of centrifugal force. But we have already shown that the ten-sions or strains upon cords, or rather the centrifugal forces, generated by the revo-ution of balls or globes, when the distan-ces are the same, are as the squares of the velocities. But the real velocity of the moon is colv, the --- th part of that beneficial results.

ence. noon is oals, the studing the part of that

mences on Monday.

as given in Herschel's Tables, in his "Out-lines of Astronomy." 14. It is a wonderful achievement in science to be able to demonstrate the law of gravity, from the simple experiment of whiring a weight, attached to a cord of riven length, around a pivet as a centre, and observing the centrifugal force gen-erated in any given time. Who could have supposed that the whiring of a ball in a semail circle of one foot radius, was a phe-nomenon of precisely the same nature, as the vevolution of magnificent worlds, in their grand orbits? and that the laws

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bodies will be inversely proportion-al to their masses; that is, if the masses of the two moving bodies are equal, their velocities towards dec., &c.

each other will be equal; if the mass of one body is twice that of the other, its velocity will be one-half; if its mass is three times greater, its velocity will be onethird; if its mass is one-hundred times greater, its velocity will be one-hundredth part of that of the

The quartity of motion in a moving body is equal to the velo-city multiplied into the mass; when two bodies move tomutual gravitation, their quantities of motion are always exactly equal. Therefore, when they meet, they will mutually destroy each other's motions, and remain at rest.

Thus it will be perceived that he force of gravitation is exhibited in two ways, namely, by pressure and by motion. Both of these ef-fects are used to measure the force. When gravity is measured by pres-sure, the result is called weight,

which a body moves by gravitation in a given time. Newton demon-strated, by careful experiments, that every kind of matter, falls from a perpendicular distance from the curve, at such straight line reaching the curve, at such straight line reaching the curve, at the end of one second, is the true measures the end of one second, is the true measures the end of one second, is the true measures the end of one second, is the true measures the end of one second. that every kind of matter near the state of rest, through equal spaces in equal times, when resistance of the atmosphere and other disturb-

ances are removed. 3.—The space through which ene body falls from a state of rest towards another, in one second of me, is one of the as measures of gravitation. The space is usually estimated in faches or fast. The space through which a body falls from rest in this latitude, is sixteen feet and one inch, or, 193 inches in one second.

4. In adopting this measure of the force of gravity, there are three iaws which should be well under-

FIRST LAW. The space through which one body, by virtue of its gravitation, moves towards another in one second, does not depend upon the mass of the former, but is the same whitever may be its mass, if the distances are the same.

Thus, the Sun gravitates towards Jupiter, and the earth gravitates towards Jupiter also; but though the mass of the Sun is 314,760 times greater than the mass of the earth, yet the gravitation of the former towards Jupiter is exactly equal to the gravitation of the latter to-wards him, when the Sun and earth are equally distant from him; in tother words, the Sun and earth, when at equal distances from Jupiter, move through the sime number of inches or parts of inches towards Jupiter in one sec. ond of time. And such must al-ways be the case, whatever may the the differences of the masses of the mass of the the mass of the sec.

ways be the case, whatever may be the differences of the masses of any number of moving bodies to-wards another body. SECOND LAW.—The space through maich confess, by their promitation, move churgeds another in one second, is proportional to the mass of the infer of The distances between the

Centrifugal Forces

6th. As we have examined the law gravitation, and determined a proper mea-sure of its force for any given quantity of ge for any given quantity of matter, and for any given distance, next, let us inquire into the nature of those op-posing forces by which the planets and satellites are prevented from falling to one ommon centre. 7.-Every mathematician is well so

f forces and motions. If a body is acted upon by two forces, it will generally have a resultant motion intermediate in direc-tion, or between the directions which would have resulted from the action of added together; make to toet, hence of dis-fail 16 feet in one second. 13.-By a combination of the law of dis-tance with the law of velocity, a law still tance with the law of pelocity, a law still

ther force singly. For instance, if a ball be projected on a smooth plane of ice, its path would be a straight line; but if the same ball be blown by a strong wind, at sime call be blown by a strong which, it right angles to its projection or path, it would deviate from a straight line lute a enrye; or, if, the ball be attached to a cord 100 feet long, and the other extremity be

fastened to a pivot, 100 feet from the straight line which the ball is describing by virtue of its projection, the ball will immediately deviate from such straight ing, or is former path, and will describe a circle around the pivot as its centre of motion. In this case, there are two forces which produce the circular motion; one is sure, the result is called weight, and is represented by the number of ounces, pounds, tons, etc. When it is measured by motion, the result is determined by the space through which a body moves by gravitation in a given time. Newton demonthe first imparts the imparts of progres-sion, and then causes to act, the second acts continually and uniformly from a central point, as steam of gravitation. The space through which the body would travel in a straight line, in one second, is a proper measure of the projectile force: the perpendicular distance from the end of

ure of the fall towards the centre. 8 —All the planets and satellites of our system are under the influence of a pro-jectile force, or a force which originally gave them motion, and then ceased to act; he same as powder ceases to act upon a will, after the moment of explosion Without a resisting medium, or some Without a resisting medium, or some other extraneous cause, the planetary badies would have continued to more for-ever, with a uniform velocity, in the straight line in which they were projected. But gravitation every moment deflects them from this right line into curvilinear orbits. The nature of these orbits will depend upon the intensity and direction of the two

9.-Ist. If the motion arising from the

9.—1st. If the motion arising from the projection is at right angles to the mo-tion arising from gravitation, and if the former is exactly proportioned to that of the latter, so as to keep the body at all times at the same distance from the centre, its orbit will be a circle of the body at all gad. Though the projectile velocity is exactly such as would produce a circular orbit, when operating at right angles to the central force, yet, if operating at an acute so obtase angle to the latter force, the body must describe an ellipse, and in this case, its mean velocity will be equal to that

must be nine times as far in the same time; for, during the first third of the time, the ply equally certain to the other? problem may be generalized thus—

Given the time of the revolution of a weight having a foot radius, in generating a cen trifugal force equal to its weight, to discove the universal law of gravitation. The method of solution has been fully de

for, during the first third of the time, the deflection will be one part of the space equal to one-nigth of the whole: the next third of the time, the body will be deflected through three parts more; and the last third of the time, it will be deflected through five parts more; one, three, and five, being added together, will be equal to nine. These are the philosophical reasons why the de-flections and forces vary directly as the square of the velocities. Falling bodies observe this same law: the first quarter of a second it fails 3 feet; the next quarter of a second it fails 3 feet; the third quarter it fails 5 feet; the fourth quarter 7 feet: one, hree, five, and seves, added together, make 16 feet; hence; bodies fail 16 feet in one second. The method of solution has been fully de-veloped in this lecture. And its truth may be verified in the most minute particulars, by its application to the grand machinery of the solarsystem. We will now give two examples. Examples.

sun on the same scale will be = 382.88and the velocity of the mean around the s will be 28.67 times its velocity around the carth, what will be the centrifugal for generated by the revolution of the mo around the sun? We have

(28.67) -- 2.147

tance with the law of velocity, a law still more general may be obtained, applicable to any given length of the strings, and to any given velocity of the whirling ball. This law may be expressed as follows:-The tensions of strings cars directly as the equirate of the velocities of the recoleing weight, attached to their ends, and inversely as their lengths.

d 282.6500 = 2.117 d 282.6500 = 2.117 Hence the centrifugal force which the mean reserves, in revolving around the sun, in company with its primary, is 2.14 times greater than that generated by he revolution around the earth. Example 2.117 the mean sun were placed at equal distances from the moon what would be the intensity of the gravita tion of the lafter towards the sun? Gravitation increases as the square of Examples:-1. A string, one foot long, with a one pound weight attached, and hav-ing a velocity of one, or which performs one complete revelution in a horizontal plane in 1.11137966 seconds, will exert a force of tension on the string equal to 1 pound.

146,600.08 × 2.147=314,760 Thus we see, that if the moon were taken as near the sun as she is to the earth, the intensity of her gravitation to the solar orb would be increased 146,000 times more than at hor present distance. But we have seen by the first example, that even in her pres-ent position, she has over twice the inten-sity of gravitation towards the sun, that abe has towards the earth; hence, the whole intensity is 314,760 times greater or in other words, the sun must contain 314,500 times more matter than the earth. If. The mechanical have which I have explained constitute a perfect balance in which to accurately determine the compar-ative weights of words. The astronomes of the present day is as familiar with the process of weighing worlds, as the chemis-is in weighing the ingredients which enter-into his compounds. These wonderful problems are now exceedingly simple, and 4 3. If the weight is attached to a cord 10 feet long, and whirled with a velocity 30 times greater than that in the first ex-ample, it will generate a centrifugal force outwards from the centre of motion, equal to 40 pounds weight; that is, the herizontal strain upon the cord will be equal to a 40-pound weight; having no motion. This is determined in the same manner as the pre- $20 \times 20$ 

is in weighing the ingredients which en-into his compounds. These wonder problems are now exceedingly simple, their solutions are within the reach of who have acquired a common school arm tion. An Arithmetic, including mecha-cal and astronomical problems of the ture of those just investigated, would be too abstrues for common school r poses, and would impire the youth of country with a love for those grand hibitions of mechanical arill, so wond-fully displayed in the mechanism Creation. 10 4. A string, 400 feet long, attached to a 1 pound weight, and having a horizontal re-volving velocity 30 times greater than that in the first example will strain the cord equal to 21 pounds;  $30 \times 30$ = 2101 4

400 Nothing is calculated to fr To determine the utmost tension that on the imparted to a string of given strength When the strength of strings is made to the aversely as the squares of their lengths, the hs greatest revolving velocities which the man with a more proto-or the Great Author of nation contemplation of his married

EXAMPLES-L A pound weight, attache

to a string of a certain given strength, one foot long, and revolving in a horizontal plane, with a velocity represented by one, will support, at its greatest possible ten-sion, before brenking, a centrifugal force represented by one. 2. A string four times as long will, ac-cording to the assumed hypothesis, he six-teen times weaker; (one-sixteenth being the inverse square of its length;) and the great-est revolving velocity which it can endure, without breaking, will be one-balf the ve-locity given in the first example, (one-half being the inverse square root of its length).

representing the central force, n times weaker, the central force, consary to balance it.

ceding; thus,

thus,-

rteen times weaker; but the h the velocity 3% will generate a c orce equal to one-sixteenth; or 1×1

3. If the string is 400 times longer which it is capable of entiring, will breaking, will be equal to one-twentie that given in the first example (one-tw ath being the inverse square root o length). In this case, the greatest per tension which the string will enture out breaking is the one funded and

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